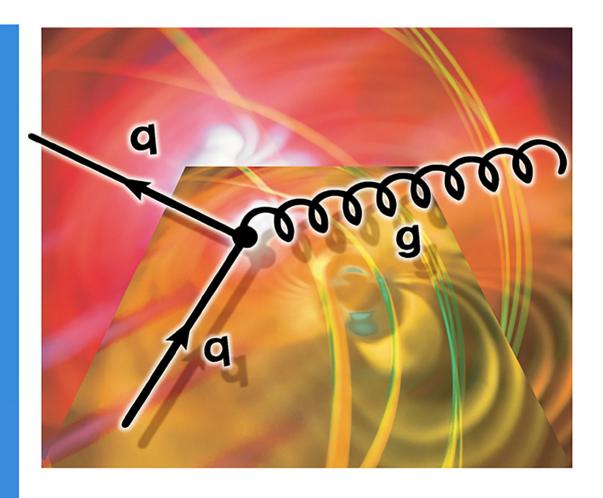
**David Griffiths** 

WILEY-VCH

# Introduction to Elementary Particles

Second, Revised Edition



David Griffiths
Introduction to Elementary
Particles

# David Griffiths

# **Introduction to Elementary Particles**

Second, Revised Edition



WILEY-VCH Verlag GmbH & Co. KGaA

#### The Author

#### David Griffiths

Reed College Portland, OR USA

For a **Solutions Manual**, instructors should contact the editorial department at vch-physics@wiley-vch.de, stating their affiliation and the course in which they wish to use the book.

1st Edition 2008
1st Reprint 2009
2nd Reprint 2010
3rd Reprint 2010
4th Reprint 2011
5th Reprint 2011
6th Reprint 2011
7th Reprint 2012
8th Reprint 2014
9th Reprint 2016

All books published by Wiley-VCH are carefully produced. Nevertheless, authors, editors, and publisher do not warrant the information contained in these books, including this book, to be free of errors. Readers are advised to keep in mind that statements, data, illustrations, procedural details or other items may inadvertently be inaccurate.

# Library of Congress Card No.:

applied for

**British Library Cataloguing-in-Publication Data** A catalogue record for this book is available from the British Library.

#### Bibliographic information published by the Deutsche Nationalbibliothek

Die Deutsche Nationalbibliothek lists this publication in the Deutsche Nationalbibliografie; detailed bibliographic data are available in the Internet at http://dnb.d-nb.de.

© 2010 WILEY-VCH Verlag GmbH & Co. KGaA, Weinheim

All rights reserved (including those of translation into other languages). No part of this book may be reproduced in any form – by photoprinting, microfilm, or any other means – nor transmitted or translated into a machine language without written permission from the publishers. Registered names, trademarks, etc. used in this book, even when not specifically marked as such, are not to be considered unprotected by law.

Typesetting Laserwords Private Limited, Chennai, India Printing Strauss GmbH, Mörlenbacg Binding Litges & Dopf Buchbinderei GmbH, Heppenheim

Printed in the Federal Republic of Germany Printed on acid-free paper

ISBN: 978-3-527-40601-2

#### **Contents**

Preface to the First Edition IXPreface to the Second Edition XIFormulas and Constants XIII

#### Introduction 1

1	Historical Introduction to the Elementary Particles 13
1.1	The Classical ERA (1897–1932) 13
1.2	The Photon (1900–1924) 15
1.3	Mesons (1934–1947) 18
1.4	Antiparticles (1930–1956) 20
1.5	Neutrinos (1930–1962) 23
1.6	Strange Particles (1947–1960) 30
1.7	The Eightfold Way (1961–1964) 35
1.8	The Quark Model (1964) 37
1.9	The November Revolution and Its Aftermath (1974–1983 and
	1995) 44
1.10	Intermediate Vector Bosons (1983) 47
1.11	The Standard Model (1978–?) 49
2	Elementary Particle Dynamics 59
2.1	The Four Forces 59
2 <b>.2</b>	Quantum Electrodynamics (QED) 60
2.3	Quantum Chromodynamics (QCD) 66
2.4	Weak Interactions 71
2.4.1	Neutral 72
2.4.2	Charged 74
2.4.2.1	e e e e e e e e e e e e e e e e e e e
2.4 <b>.2</b> .2	Quarks 75
2.4.3	Weak and Electromagnetic Couplings of W and Z 78
2.5	Decays and Conservation Laws 79
2.6	Unification Schemes 84

Introduction to Elementary Particles, Second Edition. David Griffiths Copyright © 2010 WILEY-VCH Verlag GmbH & Co. KGaA, Weinheim ISBN: 978-3-527-40601-2

VI	Contents	
•	3	Relativistic Kinematics 89
	3.1	Lorentz Transformations 89
	3.2	Four-vectors 92
	3.3	Energy and Momentum 96
	3.4	Collisions 100
	3.4.1	Classical Collisions 100
	3.4.2	Relativistic Collisions 101
	3.5	Examples and Applications 102
	4	Symmetries 115
	4.1	Symmetries, Groups, and Conservation Laws 115
	4.2	Angular Momentum 120
	4.2.1	Addition of Angular Momenta 122
	4.2.2	Spin $\frac{1}{2}$ 125
	4.3	Flavor Symmetries 129
	4.4	Discrete Symmetries 136
	4.4.1	Parity 136
	4.4.2	Charge Conjugation 142
	4.4.3	CP 144
	4.4.3.1	Neutral Kaons 145
	4.4.3.2	CP Violation 147
	4.4.4	Time Reversal and the <i>TCP</i> Theorem 149
	5	Bound States 159
	5.1	The Schrödinger Equation 159
	5.2	Hydrogen 162
	5.2.1	Fine Structure 165
	5.2.2	The Lamb Shift 166
	5.2.3	Hyperfine Splitting 167
	5.3	Positronium 169
	5.4	Quarkonium 171
	5.4.1	Charmonium 174
	5.4.2	Bottomonium 175
	5.5	Light Quark Mesons 176
	5.6	Baryons 180
	5.6.1	Baryon Wave Functions 181
	5.6.2	Magnetic Moments 189
	5.6.3	Masses 191
	6	The Feynman Calculus 197
	6.1	Decays and Scattering 197
	6.1.1	Decay Rates 197
	6.1.2	Cross Sections 199

6.2

6.2.1

The Golden Rule 203 Golden Rule for Decays 204

6.2.1.1	Two-particle Decays 206
6.2.2	Golden Rule for Scattering 208
6.2.2.1	Two-body Scattering in the CM Frame 209
6.3	Feynman Rules for a Toy Theory 211
6.3.1	Lifetime of the <i>A</i> 214
6.3.2	$A + A \rightarrow B + B$ Scattering 215
6.3.3	Higher-order Diagrams 217
7	Quantum Electrodynamics 225
7.1	The Dirac Equation 225
7.2	Solutions to the Dirac Equation 229
7.3	Bilinear Covariants 235
7.4	The Photon 238
7.5	The Feynman Rules for QED 241
7.6	Examples 245
7.7	Casimir's Trick 249
7.8	Cross Sections and Lifetimes 254
7.9	Renormalization 262
8	Electrodynamics and Chromodynamics of Quarks 275
8.1	Hadron Production in e <sup>+</sup> e <sup>-</sup> Collisions 275
8.2	Elastic Electron–Proton Scattering 279
8.3	Feynman Rules For Chromodynamics 283
8.4	Color Factors 289
8.4.1	Quark and Antiquark 289
8.4.2	Quark and Quark 292
8.5	Pair Annihilation in QCD 294
8.6	Asymptotic Freedom 298
9	Weak Interactions 307
9.1	Charged Leptonic Weak Interactions 307
9.2	Decay of the Muon 310
9.3	Decay of the Neutron 315
9.4	Decay of the Pion 321
9.5	Charged Weak Interactions of Quarks 324
9.6	Neutral Weak Interactions 329
9.7	Electroweak Unification 338
9.7.1	Chiral Fermion States 338
9.7.2	Weak Isospin and Hypercharge 342
9.7.3	Electroweak Mixing 345
10	Gauge Theories 353
10.1	Lagrangian Formulation of Classical Particle Mechanics 353
10.2	Lagrangians in Relativistic Field Theory 354
10.3	Local Gauge Invariance 358

VIII	Contents	
	10.4	Yang–Mills Theory 361
	10.5	Chromodynamics 366
	10.6	Feynman Rules 369
	10.7	The Mass Term 372
	10.8	Spontaneous Symmetry-breaking 375
	10.9	The Higgs Mechanism 378
	11	Neutrino Oscillations 387
	11.1	The Solar Neutrino Problem 387
	11.2	Oscillations 390
	11.3	Confirmation 392
	11.4	Neutrino Masses 395
	11.5	The Mixing Matrix 397
	12	Afterword: What's Next? 401
	12.1	The Higgs Boson 401
	12.2	Grand Unification 405
	12.3	Matter/Antimatter Asymmetry 409
	12.4	Supersymmetry, Strings, Extra Dimensions 411
	12.4.1	Supersymmetry 411
	12.4.2	Strings 413
	12.5	Dark Matter/Dark Energy 414
	12.5.1	Dark Matter 414
	12.5.2	Dark Energy 416
	12.6	Conclusion 417
	Α	The Dirac Delta Function 423
	В	Decay Rates and Cross Sections 429
	B.1	Decays 429
	B.1.1	Two-body Decays 429
	B.2	Cross Sections 430
	B.2.1	Two-body Scattering 430
	С	Pauli and Dirac Matrices 433
	C.1	Pauli Matrices 433
	C.2	Dirac Matrices 434
	D	Feynman Rules (Tree Level) 437
	D.1	External Lines 437
	D.2	Propagators 437
	D.3	Vertex Factors 438

Index 441

#### Preface to the First Edition

This introduction to the theory of elementary particles is intended primarily for advanced undergraduates who are majoring in physics. Most of my colleagues consider this subject inappropriate for such an audience – mathematically too sophisticated, phenomenologically too cluttered, insecure in its foundations, and uncertain in its future. Ten years ago I would have agreed. But in the last decade the dust has settled to an astonishing degree, and it is fair to say that elementary particle physics has come of age. Although we obviously have much more to learn, there now exists a coherent and unified theoretical structure that is simply too exciting and important to save for graduate school or to serve up in diluted qualitative form as a subunit of modern physics. I believe the time has come to integrate elementary particle physics into the standard undergraduate curriculum.

Unfortunately, the research literature in this field is clearly inaccessible to undergraduates, and although there are now several excellent graduate texts, these call for a strong preparation in advanced quantum mechanics, if not quantum field theory. At the other extreme, there are many fine popular books and a number of outstanding *Scientific American* articles. But very little has been written specifically for the undergraduate. This book is an effort to fill that need. It grew out of a one-semester elementary particles course I have taught from time to time at Reed College. The students typically had under their belts a semester of electromagnetism (at the level of Lorrain and Corson), a semester of quantum mechanics (at the level of Park), and a fairly strong background in special relativity.

In addition to its principal audience, I hope this book will be of use to beginning graduate students, either as a primary text, or as preparation for a more sophisticated treatment. With this in mind, and in the interest of greater completeness and flexibility, I have included more material here than one can comfortably cover in a single semester. (In my own courses I ask the students to read Chapters 1 and 2 on their own, and begin the lectures with Chapter 3. I skip Chapter 5 altogether, concentrate on Chapters 6 and 7, discuss the first two sections of Chapter 8, and then jump to Chapter 10.) To assist the reader (and the teacher) I begin each chapter with a brief indication of its purpose and content, its prerequisites, and its role in what follows.

# X | Preface to the First Edition

This book was written while I was on sabbatical at the Stanford Linear Accelerator Center, and I would like to thank Professor Sidney Drell and the other members of the Theory Group for their hospitality.

David Griffiths 1986

#### Preface to the Second Edition

It is 20 years since the first edition of this book was published, and it is both gratifying and distressing to reflect that it remains, for the most part, reasonably up-to-date. There are, to be sure, some gross lacunae - the existence of the top quark had not been confirmed back then, and neutrinos were generally assumed (for no very good reason) to be massless. But the Standard Model, which is, in essence, the subject of the book, has proved to be astonishingly robust. This is a tribute to the theory, and at the same time an indictment of our collective imagination. I don't think there has been a comparable period in the history of elementary particle physics in which so little of a truly revolutionary nature has occurred. What about neutrino oscillations? Indeed: a fantastic story (I have added a chapter on the subject); and yet, this extraordinary phenomenon fits so comfortably into the Standard Model that one might almost say, in retrospect (of course), that it would have been more surprising if it had not been so. How about supersymmetry and string theory? Yes, but these must for the moment be regarded as speculations (I have added a chapter on contemporary theoretical developments). As far as solid experimental confirmation goes, the Standard Model (with neutrino masses and mixing) still rules.

In addition to the two new chapters already mentioned, I have brought the history up-to-date in Chapter 1, shortened Chapter 5, provided (I hope) a more compelling introduction to the Golden Rule in Chapter 6, and eliminated most of the old Chapter 8 on electromagnetic form factors and scaling (this was crucially important in interpreting the deep inelastic scattering experiments that put the quark model on a secure footing, but no one today doubts the existence of quarks, and the technical details are no longer so essential). What remains of Chapter 8 is now combined with the old Chapter 9 to make a new chapter on hadrons. Finally, I have prepared a complete solution manual (available free from the publisher, though only – I regret – to course instructors). Beyond this the changes are relatively minor.

Many people have sent me suggestions and corrections, or patiently answered my questions. I cannot thank everyone, but I would like to acknowledge some of those who were especially helpful: Guy Blaylock (UMass), John Boersma (Rochester), Carola Chinellato (Brazil), Eugene Commins (Berkeley), Mimi Gerstell (Caltech),

Introduction to Elementary Particles, Second Edition. David Griffiths Copyright © 2010 WILEY-VCH Verlag GmbH & Co. KGaA, Weinheim ISBN: 978-3-527-40601-2

Nahmin Horwitz (Syracuse), Richard Kass (Ohio State), Janis McKenna (UBC), Jim Napolitano (RPI), Nic Nigro (Seattle), John Norbury (UW-Milwaukee), Jason Quinn (Notre Dame), Aaron Roodman (SLAC), Natthi Sharma (Eastern Michigan), Steve Wasserbeach (Haverford), and above all Pat Burchat (Stanford).

Part of this work was carried out while I was on sabbatical, at Stanford and SLAC, and I especially thank Patricia Burchat and Michael Peskin for making this possible.

> DAVID GRIFFITHS 2008

# Formulas and Constants Particle Data

Mass in  $MeV/c^2$ , lifetime in seconds, charge in units of the proton charge.

# Leptons (spin 1/2)

Generation	Flavor	Charge	Mass*	Lifetime	Principal Decays
first	e (electron)	-1	0.510999	$\infty$	_
	$v_e$ (e neutrino)	0	0	$\infty$	_
second	$\mu$ (muon)	-1	105.659	$2.19703 \times 10^{-6}$	$ev_{\mu}\overline{v}_{e}$
	$\nu_{\mu}$ ( $\mu$ neutrino)	0	0	$\infty$	_
third	τ (tau)	-1	1776.99	$2.91 \times 10^{-13}$	$ev_{\tau}\overline{v}_{e}, \mu v_{\tau}\overline{v}_{\mu}, \pi^{-}v_{\tau}$
	$v_{\tau}$ ( $\tau$ neutrino)	0	0	$\infty$	_

<sup>\*</sup>Neutrino masses are extremely small, and for most purposes can be taken to be zero; for details see Chapter 11.

# Quarks (spin 1/2)

Generation	Flavor	Charge	Mass*
first	d (down)	-1/3	7
	и (up)	2/3	3
second	s (strange)	-1/3	120
	c (charm)	2/3	1200
third	b (bottom)	-1/3	4300
	t (top)	2/3	174000

<sup>\*</sup>Light quark masses are imprecise and speculative; for effective masses in mesons and baryons, see Chapter 5.

# Mediators (spin 1)

Force	Mediator	Charge	Mass*	Lifetime	Principal Decays
Strong	g (8 gluons)	0	0	$\infty$	-
Electromagnetic	γ (photon)	0	0	$\infty$	-
Weak	W <sup>±</sup> (charged)	±1		$3.11 \times 10^{-25}$	$e^+\nu_e, \mu^+\nu_\mu, \tau^+\nu_\tau, cX \rightarrow \text{hadrons}$
	$Z^0$ (neutral)	0	91,190	$2.64 \times 10^{-25}$	$e^+e^-, \mu^+\mu^-, \tau^+\tau^-, q\overline{q} \rightarrow \text{hadrons}$

# Baryons (spin 1/2)

Baryon	Quark Content	Charge	Mass	Lifetime	Principal Decays
$N \left\{ p \right\}$	uud	1	938.272	$\infty$	_
n	udd	0	939.565	885.7	$pe\overline{v}_e$
Λ	uds	0	1115.68	$2.63 \times 10^{-10}$	$p\pi^-$ , $n\pi^0$
$\Sigma^+$	uus	1	1189.37	$8.02 \times 10^{-11}$	$p\pi^0$ , $n\pi^+$
$\Sigma^0$	uds	0	1192.64	$7.4 \times 10^{-20}$	$\Lambda \gamma$
$\Sigma^-$	dds	-1	1197.45	$1.48 \times 10^{-10}$	nπ <sup>-</sup>
$\Xi^0$	uss	0	1314.8	$2.90 \times 10^{-10}$	$\Lambda\pi^0$
Ξ-	dss	-1	1321.3	$1.64 \times 10^{-10}$	$\Lambda\pi^-$
$\Lambda_c^+$	udc	1	2286.5	$2.00 \times 10^{-13}$	$pK\pi$ , $\Lambda\pi\pi$ , $\Sigma\pi\pi$

# Baryons (spin 3/2)

Baryon	Quark Content	Charge	Mass	Lifetime	Principal Decays
Δ	uuu, uud, udd, ddd	2,1,0,-1	1232	$5.6 \times 10^{-24}$	Νπ
$\Sigma^*$	uus, uds, dds	1,0,-1	1385	$1.8 \times 10^{-23}$	Λπ, Σπ
$\Xi^*$	uss, dss	0,-1	1533	$6.9 \times 10^{-23}$	$\Xi\pi$
$\Omega^{-}$	SSS	-1	1672	$8.2 \times 10^{-11}$	$\Lambda K^-$ , $\Xi \pi$

# Pseudoscalar Mesons (spin 0)

Meson	Quark Content	Charge	Mass	Lifetime	Principal Decays
$\pi^\pm$	и <del>d</del> , d <del>u</del>	1,-1	139.570	$2.60 \times 10^{-8}$	$\mu \nu_{\mu}$
$\pi^0$	$(u\overline{u}-d\overline{d})/\sqrt{2}$	0	134.977	$8.4 \times 10^{-17}$	γγ
$K^{\pm}$	u <del>s</del> , s <del>u</del>	1,-1	493.68	$1.24 \times 10^{-8}$	$\mu \nu_{\mu}, \pi \pi, \pi \pi \pi$
$K^0, \overline{K}^0$	$d\overline{s}$ , $s\overline{d}$	0	497.65	$\int K_S^0 : 8.95 \times 10^{-11}$	ππ
K,K	u 3, 3 u	U	497.03	$K_L^0: 5.11 \times 10^{-8}$	$\pi e v_e$ , $\pi \mu v_\mu$ , $\pi \pi \pi$
η	$(u\overline{u} + d\overline{d} - 2s\overline{s})/\sqrt{6}$	0	547.51	$5.1 \times 10^{-19}$	γγ, πππ
$\eta'$	$(u\overline{u} + d\overline{d} + s\overline{s})/\sqrt{3}$	0	957.78	$3.2 \times 10^{-21}$	ηππ, ργ
$D^{\pm}$	$c\overline{d}$ , $d\overline{c}$	1,-1	1869.3	$1.04 \times 10^{-12}$	$Kππ$ , $Kμν_μ$ , $Keν_e$
$D^0, \overline{D}^0$	с <del>и</del> , и <del>с</del>	0	1864.5	$4.1 \times 10^{-13}$	$Kππ$ , $Keν_e$ , $Kμν_μ$
$D_{\rm s}^{\pm}$	c̄s, s̄c̄	1,-1	1968.2	$5.0 \times 10^{-13}$	$\eta \rho$ , $\phi \pi \pi$ , $\phi \rho$
$B^{\pm}$	и <del>Б</del> , b <del>u</del>	1,-1	5279.0	$1.6 \times 10^{-12}$	$D^*\ell \nu_\ell$ , $D\ell \nu_\ell$ , $D^*\pi\pi\pi$
$B^0$ , $\overline{B}^0$	$d\overline{b}, b\overline{d}$	0	5279.4	$1.5 \times 10^{-12}$	$D^*\ell v_\ell$ , $D\ell v_\ell$ , $D^*\pi\pi$

# Vector Mesons (spin 1)

Meson	Quark Content	Charge	Mass	Lifetime	Principal Decays
ρ	$u\overline{d}$ , $(u\overline{u}-d\overline{d})/\sqrt{2}$ , $d\overline{u}$	1,0,-1	775.5	$4 \times 10^{-24}$	ππ
<i>K</i> *	$u\overline{s}$ , $d\overline{s}$ , $s\overline{d}$ , $s\overline{u}$	1,0,-1	894	$1 \times 10^{-23}$	Κπ
ω	$(u\overline{u}+d\overline{d})/\sqrt{2}$	0	782.6	$8 \times 10^{-23}$	πππ, πγ
$\phi$	SS	0	1019	$2 \times 10^{-22}$	KK
$\psi$	cc	0	3097	$7 \times 10^{-21}$	$e^+e^-, \mu^+\mu^-, 5\pi, 7\pi$
$D^*$	$c\overline{d}$ , $c\overline{u}$ , $u\overline{c}$ , $d\overline{c}$	1,0,-1	2008	$3 \times 10^{-21}$	$D\pi$ , $D\gamma$
Υ	$b\overline{b}$	0	9460	$1 \times 10^{-20}$	$e^+e^-, \mu^+\mu^-, \tau^+\tau^-$

# **Spin 1/2**

Pauli Matrices:

$$\sigma_{x} \equiv \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \sigma_{y} \equiv \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \quad \sigma_{z} \equiv \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$
$$\sigma_{i}\sigma_{j} = \delta_{ij} + i\epsilon_{ijk}\sigma_{k}, \quad (\mathbf{a} \cdot \boldsymbol{\sigma})(\mathbf{b} \cdot \boldsymbol{\sigma}) = \mathbf{a} \cdot \mathbf{b} + i\boldsymbol{\sigma} \cdot (\mathbf{a} \times \mathbf{b})$$
$$\sigma_{i}^{\dagger} = \sigma_{i} = \sigma_{i}^{-1}, \qquad e^{i\boldsymbol{\theta} \cdot \boldsymbol{\sigma}} = \cos\theta + i(\hat{\theta} \cdot \boldsymbol{\sigma})\sin\theta$$

**Dirac Matrices:** 

$$\begin{split} \gamma^0 &\equiv \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}, \quad \gamma^i \equiv \begin{pmatrix} 0 & \sigma_i \\ -\sigma_i & 0 \end{pmatrix}, \quad \gamma^{0\dagger} = \gamma^0, \quad \gamma^{i\dagger} = -\gamma^i, \quad \gamma^0 \gamma^{\mu\dagger} \gamma^0 = \gamma^\mu \\ & \{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu}, \quad g^{\mu\nu} = g_{\mu\nu} = \begin{cases} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{cases} \\ & \gamma^5 \equiv i\gamma^0 \gamma^1 \gamma^2 \gamma^3 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \qquad \{\gamma^\mu, \gamma^5\} = 0, \quad (\gamma^5)^2 = 1 \end{split}$$

(For product rules and trace theorems see Appendix C.)

**Dirac Equation:** 

$$\begin{split} i\hbar\gamma^{\mu}\;\partial_{\mu}\psi - mc\,\psi &= 0 \\ (\not\!p - mc)u &= 0, \quad (\not\!p + mc)v = 0, \quad \overline{u}(\not\!p - mc) = 0, \quad \overline{v}(\not\!p + mc) = 0 \\ \overline{\psi} &\equiv \psi^{\dagger}\gamma^{0}, \quad \overline{\Gamma} \equiv \gamma^{0}\Gamma^{\dagger}\gamma^{0}, \quad \not\! A \equiv a_{\mu}\gamma^{\mu} \end{split}$$

# Feynman Rules

	External Lines		Propagators	
Spin 0:	Nothing	$\frac{i}{q^2 - (mc)^2}$		
	Incoming particle:	и		
Spin 1/2:	Incoming particle: Incoming antiparticle: Outgoing particle: Outgoing antiparticle:	$\overline{v}$	i(q + mc)	
Spiii 1/2.	Outgoing particle:	ū	$\frac{i(q+mc)}{q^2-(mc)^2}$	
	Outgoing antiparticle:	v		

Spin 1: 
$$\begin{cases} \text{Incoming: } \epsilon_{\mu} \\ \text{Outgoing: } \epsilon_{\mu}^{*} \end{cases} \begin{cases} \text{Massless: } \frac{-ig_{\mu\nu}}{q^{2}} \\ \text{Massive: } \frac{-i[g_{\mu\nu} - q_{\mu}q_{\nu}/(mc)^{2}]}{q^{2} - (mc)^{2}} \end{cases}$$

(For vertex factors see Appendix D.)

### **Fundamental Constants**

Planck's constant:  $\hbar = 1.05457 \times 10^{-34} \text{ J s}$ 

 $= 6.58212 \times 10^{-22} \text{ MeV s}$ 

Speed of light:  $c = 2.99792 \times 10^8 \text{ m/s}$ 

Mass of electron:  $m_e = 9.10938 \times 10^{-31} \text{kg} = 0.510999 \text{ MeV/c}^2$ 

Mass of proton:  $m_p = 1.67262 \times 10^{-27} \text{kg} = 938.272 \text{ MeV/c}^2$ 

Electron charge (magnitude):  $e = 1.60218 \times 10^{-19} \text{ C}$ 

 $= 4.80320 \times 10^{-10} \text{ esu}$ 

Fine structure constant:  $\alpha = e^2/\hbar c = 1/137.036$ 

Bohr radius:  $a = \hbar^2/m_e e^2 = 5.29177 \times 10^{-11} \text{ m}$ 

Bohr energies:  $E_n = -m_e e^4/2\hbar^2 n^2 = -13.6057/n^2 \text{ eV}$ 

Classical electron radius:  $r_e = e^2/m_e c^2 = 2.81794 \times 10^{-15} \text{ m}$ 

QED coupling constant:  $g_e = e\sqrt{4\pi/\hbar c} = 0.302822$ 

Weak coupling constants:  $g_w = g_e / \sin \theta_w = 0.6295;$ 

 $g_z = g_w/\cos\theta_w = 0.7180$ 

Weak mixing angle:  $\theta_w = 28.76^{\circ} (\sin^2 \theta_w = 0.2314)$ 

Strong coupling constant:  $g_s = 1.214$ 

## **Conversion Factors**

1 Å =  $0.1 \text{ nm} = 10^{-10} \text{ m}$ 

1 fm =  $10^{-15}$  m 1 barn =  $10^{-28}$  m<sup>2</sup>

1 eV =  $1.60218 \times 10^{-19} \text{ J}$ 

 $1 \text{ MeV}/c^2 = 1.78266 \times 10^{-30} \text{ kg}$ 

1 Coulomb =  $2.99792 \times 10^9 \text{ esu}$ 

#### Introduction

#### **Elementary Particle Physics**

Elementary particle physics addresses the question, 'What is matter made of?' at the most fundamental level – which is to say, on the smallest scale of size. It's a remarkable fact that matter at the subatomic level consists of tiny chunks, with vast empty spaces in between. Even more remarkable, these tiny chunks come in a small number of different types (electrons, protons, neutrons, pi mesons, neutrinos, and so on), which are then replicated in astronomical quantities to make all the 'stuff' around us. And these replicas are absolutely perfect copies – not just 'pretty similar', like two Fords coming off the same assembly line, but utterly *indistinguishable*. You can't stamp an identification number on an electron, or paint a spot on it – if you've seen one, you've seen them all. This quality of absolute identicalness has no analog in the macroscopic world. (In quantum mechanics it is reflected in the Pauli exclusion principle.) It enormously simplifies the task of elementary particle physics: we don't have to worry about big electrons and little ones, or new electrons and old ones – an electron is an electron is an electron. It didn't have to be so easy.

My first job, then, is to introduce you to the various kinds of elementary particles – the actors, if you will, in the drama. I could simply *list* them, and tell you their properties (mass, electric charge, spin, etc.), but I think it is better in this case to adopt a historical perspective, and explain how each particle first came on the scene. This will serve to endow them with character and personality, making them easier to remember and more interesting to watch. Moreover, some of the stories are delightful in their own right.

Once the particles have been introduced, in Chapter 1, the issue becomes, 'How do they interact with one another?' This question, directly or indirectly, will occupy us for the rest of the book. If you were dealing with two *macroscopic* objects, and you wanted to know how they interact, you would probably begin by holding them at various separation distances and measuring the force between them. That's how Coulomb determined the law of electrical repulsion between two charged pith balls, and how Cavendish measured the gravitational attraction of two lead weights. But you can't pick up a proton with tweezers or tie an electron onto the end of a piece of string; they're just too small. For practical reasons, therefore, we have to resort to

Introduction to Elementary Particles, Second Edition. David Griffiths Copyright © 2010 WILEY-VCH Verlag GmbH & Co. KGaA, Weinheim ISBN: 978-3-527-40601-2

less direct means to probe the interactions of elementary particles. As it turns out, almost all of our experimental information comes from three sources: (1) scattering events, in which we fire one particle at another and record (for instance) the angle of deflection; (2) decays, in which a particle spontaneously disintegrates and we examine the debris; and (3) bound states, in which two or more particles stick together, and we study the properties of the composite object. Needless to say, determining the interaction law from such indirect evidence is not a trivial task. Ordinarily, the procedure is to guess a form for the interaction and compare the resulting theoretical predictions with the experimental data.

The formulation of such a guess ('model' is a more respectable term for it) is guided by certain general principles, in particular, special relativity and quantum mechanics. In the diagram below I have sketched out four realms of mechanics:

	$Small \to$				
	Classical	Quantum			
<b>.</b>	mechanics	mechanics			
Fast ↓	Relativistic mechanics	Quantum field theory			

The world of everyday life, of course, is governed by classical mechanics. But for objects that travel very fast (at speeds comparable to c), the classical rules are modified by special relativity, and for objects that are very small (comparable to the size of atoms, roughly speaking), classical mechanics is superseded by quantum mechanics. Finally, for things that are both fast and small, we require a theory that incorporates relativity and quantum principles: quantum field theory. Now, elementary particles are extremely small, of course, and typically they are also very fast. So, elementary particle physics naturally falls under the dominion of quantum field theory.

Please observe the distinction here between a type of mechanics and a particular force law. Newton's law of universal gravitation, for example, describes a specific interaction (gravity), whereas Newton's three laws of motion define a mechanical system (classical mechanics), which (within its jurisdiction) governs all interactions. The force law tells you what *F* is, in the case at hand; the mechanics tells you how to use F to determine the motion. The goal of elementary particle dynamics, then, is to guess a set of force laws which, within the context of quantum field theory, correctly describe particle behavior.

However, some general features of this behavior have nothing to do with the detailed form of the interactions. Instead they follow directly from relativity, from quantum mechanics, or from the combination of the two. For example, in relativity, energy and momentum are always conserved, but (rest) mass is not. Thus the decay  $\Delta \to p + \pi$  is perfectly acceptable, even though the  $\Delta$  weighs more than the sum of p plus  $\pi$ . Such a process would not be possible in classical mechanics, where mass is strictly conserved. Moreover, relativity allows for particles of zero (rest) mass – the very idea of a massless particle is nonsense in classical mechanics – and as we shall see, photons and gluons are massless.

In quantum mechanics a physical system is described by its state, s (represented by the wave function  $\psi_s$  in Schrödinger's formulation, or by the ket  $|s\rangle$  in Dirac's theory). A physical process, such as scattering or decay, consists of a transition from one state to another. But in quantum mechanics the outcome is not uniquely determined by the initial conditions; all we can hope to calculate, in general, is the probability for a given transition to occur. This indeterminacy is reflected in the observed behavior of particles. For example, the charged pi meson ordinarily disintegrates into a muon plus a neutrino, but occasionally one will decay into an *electron* plus a neutrino. There's no difference in the original pi mesons; they're all identical. It is simply a fact of nature that a given particle can go either way.

Finally, the union of relativity and quantum mechanics brings certain extra dividends that neither one can offer by itself: the existence of antiparticles (with the same mass and lifetime as the particle itself, but opposite electric charge), a proof of the Pauli exclusion principle (which in nonrelativistic quantum mechanics is simply an ad hoc hypothesis), and the so-called TCP theorem. I'll tell you more about these later on; my purpose in mentioning them here is to emphasize that these are features of the mechanical system itself, not of the particular model. Short of a catastrophic revolution, they are untouchable. By the way, quantum field theory in all its glory is difficult and deep, but don't be alarmed: Feynman invented a beautiful and intuitively satisfying formulation that is not hard to learn; we'll come to that in Chapter 6. (The derivation of Feynman's rules from the underlying quantum field theory is a different matter, which can easily consume the better part of an advanced graduate course, but this need not concern us here.)

In the 1960s and 1970s a theory emerged that described all of the known elementary particle interactions, except gravity. (As far as we can tell, gravity is much too weak to play any significant role in ordinary particle processes.) This theory - or, more accurately, this collection of related theories, based on two families of elementary particles (quarks and leptons), and incorporating quantum electrodynamics, the Glashow-Weinberg-Salam theory of electroweak processes, and quantum chromodynamics - has come to be called the Standard Model. No one pretends that it is the final word on the subject, but at least we are now playing with a full deck of cards. Since 1978, when the Standard Model achieved the status of 'orthodoxy', it has met every experimental test. Moreover, it has an attractive aesthetic feature: all of the fundamental interactions derive from one general principle, the requirement of local gauge invariance. It seems certain that future developments will involve extensions of the Standard Model, not its repudiation. This book might be called an 'Introduction to the Standard Model'.

As that alternative title suggests, it is a book about elementary particle theory, with very little on experimental methods or instrumentation. These are important matters, and an argument can be made for integrating them into a text such as this, but they can also be distracting, and interfere with the clarity and elegance of the theory itself. I encourage you to read about the experimental aspects of the subject, and from time to time I will refer you to particularly accessible accounts. But for now, I'll confine myself to scandalously brief answers to the two most obvious experimental questions.

#### How Do You Produce Elementary Particles?

Electrons and protons are no problem; these are the stable constituents of ordinary matter. To produce electrons one simply heats up a piece of metal, and they come boiling off. If you want a *beam* of electrons, you just set up a positively charged plate nearby, to attract them over, and poke a small hole in it; the electrons that make it through the hole constitute the beam. Such an *electron gun* is the starting element in a television tube or an oscilloscope or an electron accelerator (Figure I.1).

To obtain protons you ionize hydrogen (in other words, strip off the electron). In fact, if you're using the protons as a *target*, you don't even need to bother about the electrons; they're so light that an energetic incident particle will knock them out of the way. Thus, a tank of hydrogen *is* essentially a tank of protons. For more exotic particles there are three main sources: cosmic rays, nuclear reactors, and particle accelerators.

• *Cosmic rays:* The earth is constantly bombarded with high-energy particles (principally protons) coming from outer space. What the source of these particles might be remains something of a mystery; at any rate, when they hit atoms in the upper atmosphere they produce showers of secondary particles (mostly muons and neutrinos, by the

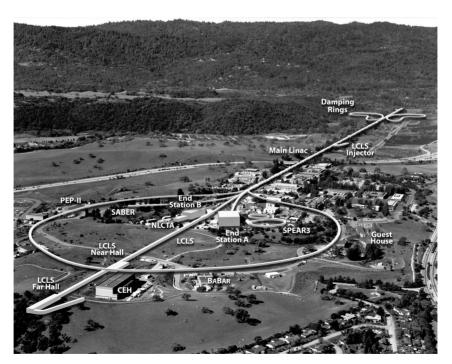


Fig. 1.1 SLAC; the straight line is the accelerator itself. (Courtesy Stanford Linear Accelerator Center.)

time they reach ground level), which rain down on us all the time. As a source of elementary particles, cosmic rays have two virtues: they are free, and their energies can be enormous – far greater than we could possibly produce in the laboratory. But they have two major disadvantages: the rate at which they strike any detector of reasonable size is very low, and they are completely uncontrollable. So cosmic ray experiments call for patience and luck.

- Nuclear reactors: When a radioactive nucleus disintegrates, it may emit a variety of particles - neutrons, neutrinos, and what used to be called alpha rays (actually, alpha particles, which are bound states of two neutrons plus two protons), beta rays (actually, electrons or positrons), and gamma rays (actually, photons).
- Particle accelerators: You start with electrons or protons, accelerate them to high energy, and smash them into a target (Figure I.1). By skillful arrangements of absorbers and magnets, you can separate the particle species that you wish to study from the resulting debris. Nowadays it is possible in this way to generate intense secondary beams of positrons, muons, pions, kaons, *B*-mesons, antiprotons, and neutrinos, which in turn can be fired at another target. The stable particles - electrons, protons, positrons, and antiprotons can even be fed into giant storage rings in which, guided by powerful magnets, they circulate at high speed for hours at a time, to be extracted and used at the required moment [1].

In general, the heavier the particle you want to produce, the higher must be the energy of the collision. That's why, historically, lightweight particles tend to be discovered first, and as time goes on, and accelerators become more powerful, heavier and heavier particles are found. It turns out that you gain enormously in relative energy if you collide two high-speed particles head-on, as opposed to firing one particle at a stationary target. (Of course, this calls for much better aim!) For this reason many contemporary experiments involve colliding beams from intersecting storage rings; if the particles miss on the first pass, they can try again the next time around. Indeed, with electrons and positrons (or protons and antiprotons) the same ring can be used, with the plus charges circulating in one direction and minus charges the other. Unfortunately, when a charged particle accelerates it radiates, thereby losing energy. In the case of circular motion (which, of course, involves acceleration) this is called *synchrotron radiation*, and it severely limits the efficiency of storage rings for energetic electrons (heavier particles with the same energy accelerate less, so synchrotron radiation is not such a problem for them). For this reason electron scattering experiments will increasingly turn to linear colliders, while storage rings will continue to be used for protons and heavier particles.

There is another reason why particle physicists are always pushing for higher energies: in general, the higher the energy of the collision, the closer the two particles come to one another. So if you want to study an interaction at very short range, you need very energetic particles. In quantum-mechanical terms, a particle of momentum p has an associated wavelength  $\lambda$  given by the de Broglie formula  $\lambda = h/p$ , where h is Planck's constant. At large wavelengths (low momenta) you can only hope to resolve relatively large structures; in order to examine something extremely small, you need comparably short wavelengths, and hence high momenta. If you like, consider this a manifestation of the uncertainty principle  $(\Delta x \Delta p \geq h/4\pi)$  – to make  $\Delta x$  small,  $\Delta p$  must be large. However you look at it, the conclusion is the same: to probe small distances you need high energies.

At present the most powerful accelerator in the world is the *Tevatron* at Fermilab (Figure I.2), with a maximum beam energy of almost 1 TeV. The tevatron (a proton—antiproton collider) began operation in 1983; its successor, the Superconducting Supercollider (SSC) was under construction in 1993 when the project was terminated by Congress. As a result, there has been a long period in which no fundamental progress was possible. This dry spell should end in 2008, when the Large Hadron Collider (LHC) at CERN starts taking data (Figure I.3). The LHC is designed to reach beam energies in excess of 7 TeV, and the hope is that this new terrain will include the Higgs particle, possibly supersymmetric particles, and — best of all — something completely unexpected [2]. It's not clear what comes after the LHC — most likely the proposed International Linear Collider (ILC). But, accelerators have become so huge (the SSC would have been 87 km in circumference) that there is not much room for expansion. Perhaps we are approaching the end



Fig. 1.2 Fermilab; the large circle in the background is the Tevatron. (Courtesy Fermilab Visual Media Services.)



Fig. 1.3 CERN; the circle indicates the path of the LHC tunnel (formerly LEP) - Geneva and Mt Blanc are in the background. (Courtesy CERN.)

of the accelerator era, and particle physicists will have to turn to astrophysics and cosmology for information about higher energies. Or perhaps someone will have a clever new idea for squeezing energy onto an elementary particle.\*

#### How Do You Detect Elementary Particles?

There are many kinds of particle detectors – Geiger counters, cloud chambers, bubble chambers, spark chambers, drift chambers, photographic emulsions, Čerenkov counters, scintillators, photomultipliers, and so on. Actually, a typical modern

<sup>\*</sup> In macroscopic terms the amount of energy involved is not that great - after all, 1 TeV  $(10^{12} \text{ eV})$  is only  $10^{-7}$  Joules; the problem is how to deliver that energy to a particle. No law of physics prevents you from doing so, but nobody has yet figured out a way to do it without gigantic (and expensive) machinery.

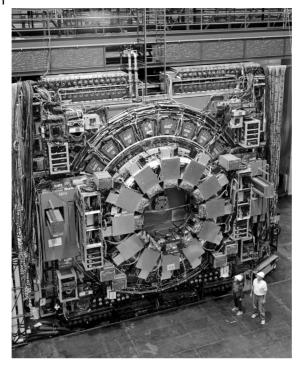


Fig. 1.4 The CDF detector at Fermilab, where the top quark was discovered. (Courtesy Fermilab Visual Media Services.)

detector has whole arrays of these devices, wired up to a computer that tracks the particles and displays their trajectories on a television screen (Figure I.4). The details do not concern us, but there is one thing you should be aware of: most detection mechanisms rely on the fact that when high-energy charged particles pass through matter they ionize atoms along their path. The ions then act as 'seeds' in the formation of droplets (cloud chamber) or bubbles (bubble chamber) or sparks (spark chamber), as the case may be. But electrically neutral particles do not cause ionization, and they leave no tracks. For instance, if you look at the bubble chamber photograph in Figure 1.9, you will see that the five neutral particles are 'invisible'; their paths have been reconstructed by analyzing the tracks of the charged particles in the picture and invoking conservation of energy and momentum at each vertex. Notice also that most of the tracks in the picture are curved (actually, all of them are, to some extent; try holding a ruler up to one you think is straight). The bubble chamber was placed between the poles of a giant magnet; in a magnetic field B, a particle of charge q and momentum p will move in a circle of radius R given by the famous *cyclotron formula*: R = pc/qB, where *c* is the speed of light. The curvature of the track in a known magnetic field thus affords a very simple measure of the particle's momentum. Moreover, we can immediately tell the sign of the charge from the direction of the curve.

Elementary particles are small, so for our purposes the normal mechanical units – grams, ergs, joules, and so on - are inconveniently large. Atomic physicists introduced the *electron volt* – the energy acquired by an electron when accelerated through a potential difference of 1 volt:  $1 \text{ eV} = 1.6 \times 10^{-19}$  joules. For us the eV is inconveniently *small*, but we're stuck with it. Nuclear physicists use keV (10<sup>3</sup> eV); typical energies in particle physics are MeV (10<sup>6</sup> eV), GeV (10<sup>9</sup> eV), or even TeV ( $10^{12}$  eV). Momenta are measured in MeV/c (or GeV/c, or whatever), and masses in MeV/ $c^2$ . Thus the proton weighs 938 MeV/ $c^2 = 1.67 \times 10^{-24}$  g.

Actually, particle theorists are lazy (or clever, depending on your point of view) – they seldom include the c's and  $\hbar$ 's ( $\hbar \equiv h/2\pi$ ) in their formulas. You're just supposed to fit them in for yourself at the end, to make the dimensions come out right. As they say in the business, 'set  $c = \hbar = 1$ '. This amounts to working in units such that time is measured in centimeters and mass and energy in inverse centimeters; the unit of time is the time it takes light to travel 1 cm, and the unit of energy is the energy of a photon whose wavelength is  $2\pi$  cm. Only at the end of the problem do we revert to conventional units. This makes everything look very elegant, but I thought it would be wiser in this book to keep all the c's and  $\hbar$ 's where they belong, so that you can check for dimensional consistency as you go along. (If this offends you, remember that it is easier for you to ignore an  $\hbar$  you don't like than for someone else to conjure one up in just the right place.)

Finally, there is the question of what units to use for electric charge. In introductory physics courses most instructors favor the SI system, in which charge is measured in coulombs, and Coulomb's law reads

$$F = \frac{1}{4\pi\epsilon_0} \frac{q_1 q_2}{r^2} \qquad (SI)$$

Most advanced work is done in the Gaussian system, in which charge is measured in electrostatic units (esu), and Coulomb's law is written

$$F = \frac{q_1 q_2}{r^2} \qquad (G)$$

But elementary particle physicists prefer the Heaviside - Lorentz system, in which Coulomb's law takes the form

$$F = \frac{1}{4\pi} \frac{q_1 q_2}{r^2}$$
 (HL)

The three units of charge are related as follows:

$$q_{\rm HL} = \sqrt{4\pi} \, q_{\rm G} = \frac{1}{\sqrt{\epsilon_0}} \, q_{\rm SI}$$

In this book I shall use Gaussian units exclusively, in order to avoid unnecessary confusion in an already difficult subject. Whenever possible I will express results in terms of the fine structure constant

$$\alpha = \frac{e^2}{\hbar c} = \frac{1}{137.036}$$

where e is the charge of the electron in Gaussian units. Most elementary particle texts write this as  $e^2/4\pi$ , because they are measuring charge in Heaviside–Lorentz units and setting  $c = \hbar = 1$ ; but everyone agrees that the *number* is 1/137.

#### Further reading

Since the early 1960s, the Particle Data Group at Berkeley has periodically issued a listing of the established particles and their properties. These are published every other year in Reviews of Modern Physics or Journal of Physics G, and summarized in a (free) booklet that can be ordered on the web at http://pdg.lbl.gov. In the early days this summary took the form of 'wallet cards', but by 2006 it had grown to a densely packed 315 pages. I shall refer to it as the Particle Physics Booklet (PPB). Every student of elementary particle physics must have a copy – don't leave home without it! The longer version, called the Review of Particle Physics (RPP) is the bible for professionals – the 2006 edition runs to 1231 pages, and it includes authoritative articles on every relevant subject, written by the world's leading experts [3]. If you want the definitive, up-to-date word on any particular topic, this is the place to go (it is also available on-line, at the Particle Data Group web site).

Particle physics is an enormous and rapidly changing subject. My aim in this book is to introduce you to some important ideas and methods, to give you a sense of what's out there to be learned, and perhaps to stimulate your appetite for more. If you want to read further in quantum field theory, I particularly recommend:

Bjorken, J. D. and Drell, S. D. (1964) Relativistic Quantum Mechanics and Relativistic Quantum Fields, McGraw-Hill, New York.

Itzykson, C. and Zuber, J.-B. (1980) Quantum Field Theory, McGraw-Hill, New York.

Peskin, M. E. and Schroeder, D. V. (1995) An Introduction to Quantum Field Theory, Perseus, Cambridge, MA.

Ryder, L. H. (1985) Quantum Field Theory, Cambridge University Press, Cambridge, UK.

Sakurai, J. J. (1967) Advanced Quantum Mechanics, Addison-Wesley, Reading, MA.

I warn you, however, that these are all difficult and advanced books.

For elementary particle physics itself, the following books (listed in order of increasing difficulty) are especially useful:

- Close, F., Marten, M. and Sutton, C. (1987) The Particle Explosion, Oxford University Press, Oxford, UK.
- Frauenfelder, H. and Henley, E. M. (1991) Subatomic Physics, 2nd edn, Prentice-Hall, Englewood Cliffs, NJ.
- Gottfried, K. and Weisskopf, V. F. (1984) Concepts of Particle Physics, Oxford University Press, Oxford.
- Perkins, D. H. (2000) Introduction to High-Energy Physics, 4th Ed, Cambridge University Press, Cambridge, UK.
- Halzen, F. and Martin, A. D. (1984) Quarks and Leptons, John Wiley & Sons, Ltd. New York.
- Roe, B. P. (1996) Particle Physics at the New Millennium, Springer, New York.
- Aitchison, I. J. R. and Hey, A. J. G. (2003) Gauge Theories in Particle Physics, 3rd edn, Institute of Physics, Bristol, UK.
- Seiden, A. (2005) Particle Physics: A Comprehensive Introduction, Addison-Wesley, San Francisco, CA.
- Quigg, C. (1997) Gauge Theories of the Strong, Weak, and Electromagnetic Interactions, Addison-Wesley, Reading, MA.

#### References

- 1 For a comprehensive bibliography see Chao, Q. W. (2006) American Journal of Physics, 74, 855.
- 2 Smith, C. L. (July 2000) Scientific American, 71; (a) Lederman, L. (2007) Nature, 448, 310.
- 3 The current reference is Yao, W.-M. et al. (2006) Journal of Physics, G 33,
- 1. But because you will want to use the most recent versions, I will simply refer to them as Particle Physics Booklet and Review of Particle Physics, appending the year when it is relevant.